

Problem 2.16

Assume;

1. steady - state
2. $\delta/R \ll 1$, so that a planar, 2-D stability analysis at the interphase is justified
3. $M_G \ll M_L$, so that the gas can be treated as inviscid
4. Axis-symmetric flow
5. Negligible gravitational effect

2.16.2

The conservation equations for the liquid phase can then be represented as:

$$\frac{\partial u_L}{\partial x} + \frac{\partial v_L}{\partial y} = 0 \quad (5)$$

$$\rho_L \frac{\partial u_L}{\partial t} = -\frac{\partial P_L}{\partial x} + \mu_L \left(\frac{\partial^2 u_L}{\partial x^2} + \frac{\partial^2 u_L}{\partial y^2} \right) \quad (6)$$

$$\rho_L \frac{\partial v_L}{\partial t} = -\frac{\partial P_L}{\partial y} + \mu_L \left(\frac{\partial^2 v_L}{\partial x^2} + \frac{\partial^2 v_L}{\partial y^2} \right) \quad (7)$$

The gas phase conservation equations are similar, except that they do not include viscosity.

Periodic motion of incompressible fluid surfaces obeying the Navier-Stokes equations can be represented using a potential and a stream function. Each velocity component of the liquid is therefore assumed to consist of an inviscid term and a perturbation representing the effect of viscosity [17].

$$u_L = u_L^0 + U_L \quad (8)$$

$$v_L = v_L^0 + V_L \quad (9)$$

$$P = P^0 \quad (10)$$

where parameters with superscript 0 represent ideal fluid conditions, and U and V represent the velocities in x and y directions due to viscosity, respectively. The gas phase is assumed to be inviscid; the gas velocity components therefore do not include terms representing viscosity. The above ideal-fluid velocities are represented by a potential function of the form:

$$\Phi_L = A_L e^{-ky + \omega t} \cos kx \quad (11)$$

$$\Phi_G = A_G \cosh[k(y + \delta)] e^{\omega t} \cos kx \quad (12)$$

where for both phases:

$$(u^o, v^o) = \left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y} \right) \Phi \quad (13)$$

These potential functions fully represent the irrotational, ideal liquid and gas fields. The liquid velocity terms that account for the viscosity-induced rotational motion can be represented by a stream function, ψ_L , that must satisfy [17]:

$$\frac{\partial \psi_L}{\partial t} = \nu_L \nabla^2 \psi_L \quad (14)$$

where:

$$(U_L, V_L) = \left(-\frac{\partial}{\partial y}, \frac{\partial}{\partial x} \right) \psi_L \quad (15)$$

Equation (14) and all the necessary boundary conditions can be satisfied using:

$$\psi_L = B_L e^{-my + \omega t} \sin kx \quad (16)$$

Note that the above equation approximates the liquid core as infinitely-deep. Using Eq. (14), it can be shown that:

$$m^2 = k^2 + \frac{\omega}{\nu_L} \quad (17)$$

The above potential and stream functions are now utilized in a linear stability analysis, using the following boundary conditions at the interphase:

$$v_L = v_G \quad (18)$$

$$-P_L + 2\mu_L \left(\frac{\partial V_L}{\partial y} \right) = -P_G - \sigma \frac{\partial^2 \eta}{\partial x^2} - \Delta P_\sigma \quad (19)$$

$$\mu_L \left(\frac{\partial u_L}{\partial y} + \frac{\partial v_L}{\partial x} \right) = 0 \quad (20)$$

where:

$$\eta = \frac{k}{\omega} (B_L - A_L) e^{\omega t} \cos kx \quad (21)$$

$$P_L - P_G = -\rho_L \left. \frac{\partial \Phi_L}{\partial t} \right|_{y=0} + \rho_G \left. \frac{\partial \Phi_G}{\partial t} \right|_{y=0} \quad (22)$$

The term ΔP_σ in Eqn. (19), which is de-stabilizing, is the surface tension force caused by the curvature of the channel, and can be represented as:

$$\Delta P_\sigma = \frac{\sigma}{R-\delta} - \frac{\sigma}{R-\delta-\eta} \approx \frac{\sigma}{R-\delta} \left(1 + \frac{\eta}{R-\delta} \right) \quad (23)$$

Equations (19-21) following substitution for velocities using the aforementioned potential and stream functions, lead to:

$$-A_L + B_L - A_G \sinh(kh) = 0 \quad (24)$$

$$2k^2 A_L - (m^2 + k^2) B_L = 0 \quad (25)$$

$$\begin{aligned} & \left[2\mu_L k^2 + \frac{k^3 \sigma}{\omega} - \frac{\sigma}{(R-\delta)^2} \frac{k}{\omega} + \rho_L \omega \right] A_L + \\ & \left[-2\mu_L km - \frac{k^3 \sigma}{\omega} + \frac{\sigma}{(R-\delta)^2} \frac{k}{\omega} \right] B_L - [\rho_G \alpha \cosh(kh)] A_G = 0 \end{aligned} \quad (26)$$

A non-trivial solution is possible if the determinant of the coefficient matrix of Eqns. (24-26) vanishes, and that leads to:

$$\left[\rho^* \coth(R^*KH) + 1 \right] \Omega^2 - \frac{K}{R^{*2}(1-H)^2} + K^3 + 4K^4 \left[1 - \left(1 + \frac{\Omega}{K^2} \right)^{1/2} \right] + 4\Omega K^2 = 0 \quad (27)$$

where $\rho^* = \rho_G / \rho_L$; $H = \delta / R$; $R^* = R \frac{\sigma}{\rho_L v_L^2}$; $K = k \left(\frac{\sigma}{\rho_L v_L^2} \right)^{-1}$; and $\Omega = \omega \left(\frac{\rho_L^2 v_L^3}{\sigma^2} \right)$. The

interface is evidently unstable when $\Omega > 0$, and the neutral condition ($\Omega = 0$) occurs when

$$(KR^*)_{cr} = \frac{1}{1-H} \quad (28)$$

Disturbances with wavelengths longer than λ_{cr} thus cause instability, where:

$$\lambda_{cr} = 2\pi(R - \delta) \quad (29)$$

The neutral wavelength thus is similar to the predictions of the Kelvin-Helmholtz stability; and does not depend on liquid viscosity, in agreement with the Taylor stability analysis for film boiling of viscous liquids. The fastest growing wavelength, Ω_d , occurs when $d\Omega/dK = 0$, whereby:

$$\left(\frac{\rho^* R^* H}{\sinh^2(R^*KH)} \right) \Omega_d^2 - 4K \left[2 + \left(1 + \frac{\Omega_d}{K^2} \right)^{-1/2} \right] \Omega_d - 16K_d^3 \left[1 - \left(1 + \frac{\Omega_d}{K_d^2} \right)^{1/2} \right] - 3K_d^2 = 0 \quad (30)$$